

# The Physics of Pulsar Magnetospheres

## June 6th 2016 - NASA GSFC

*Developing of a specialized Particle-In-Cell code to study  
Pulsar Magnetospheres*

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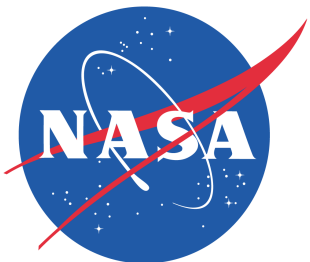
I'm a grad student and I work with:

Alice Harding

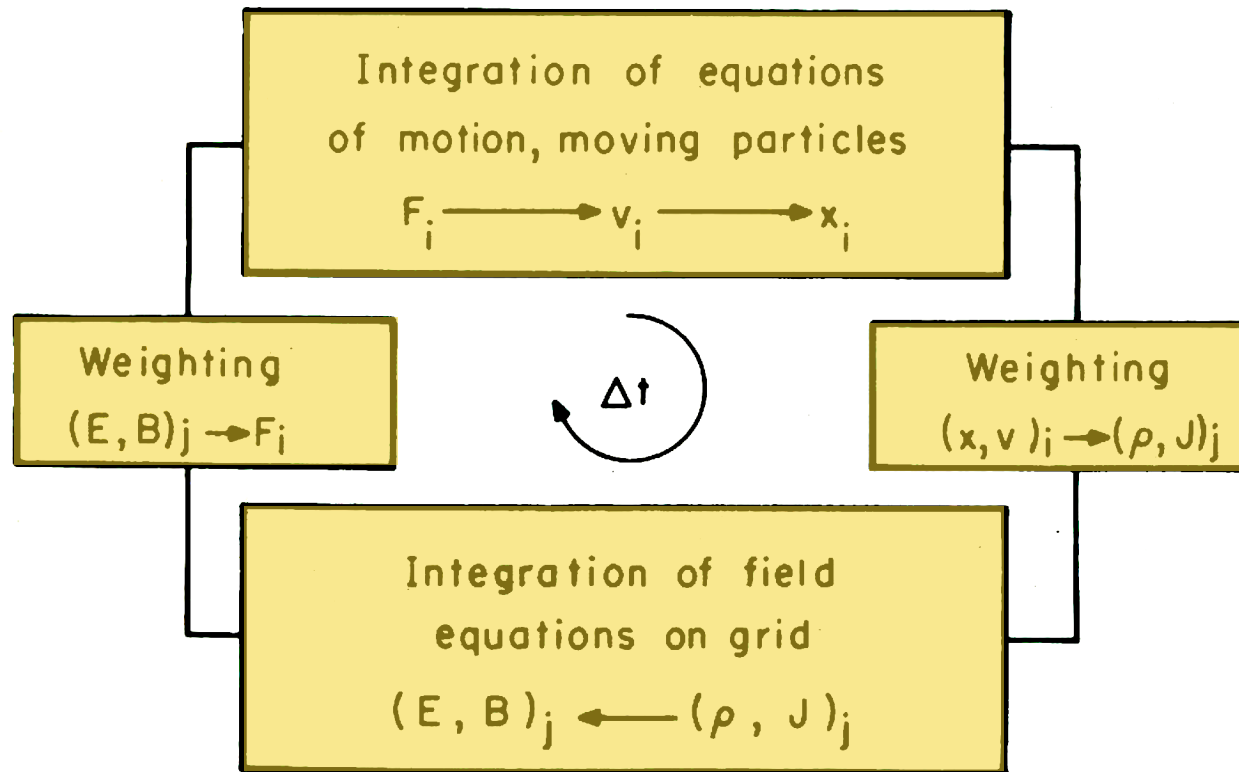
Andrey Timokhin

Constantinos Kalapotharakos

Demosthenes Kazanas



# In PIC codes, particles moved by the fields form the currents that act on the fields themselves



Birdsall & Langdon 1985 *Plasma Physics via Computer Simulation* (New York: McGraw-Hill)

3D Cartesian  
Electromagnetic  
Relativistic pusher  
Parallelized

PML  
perfectly matched layer

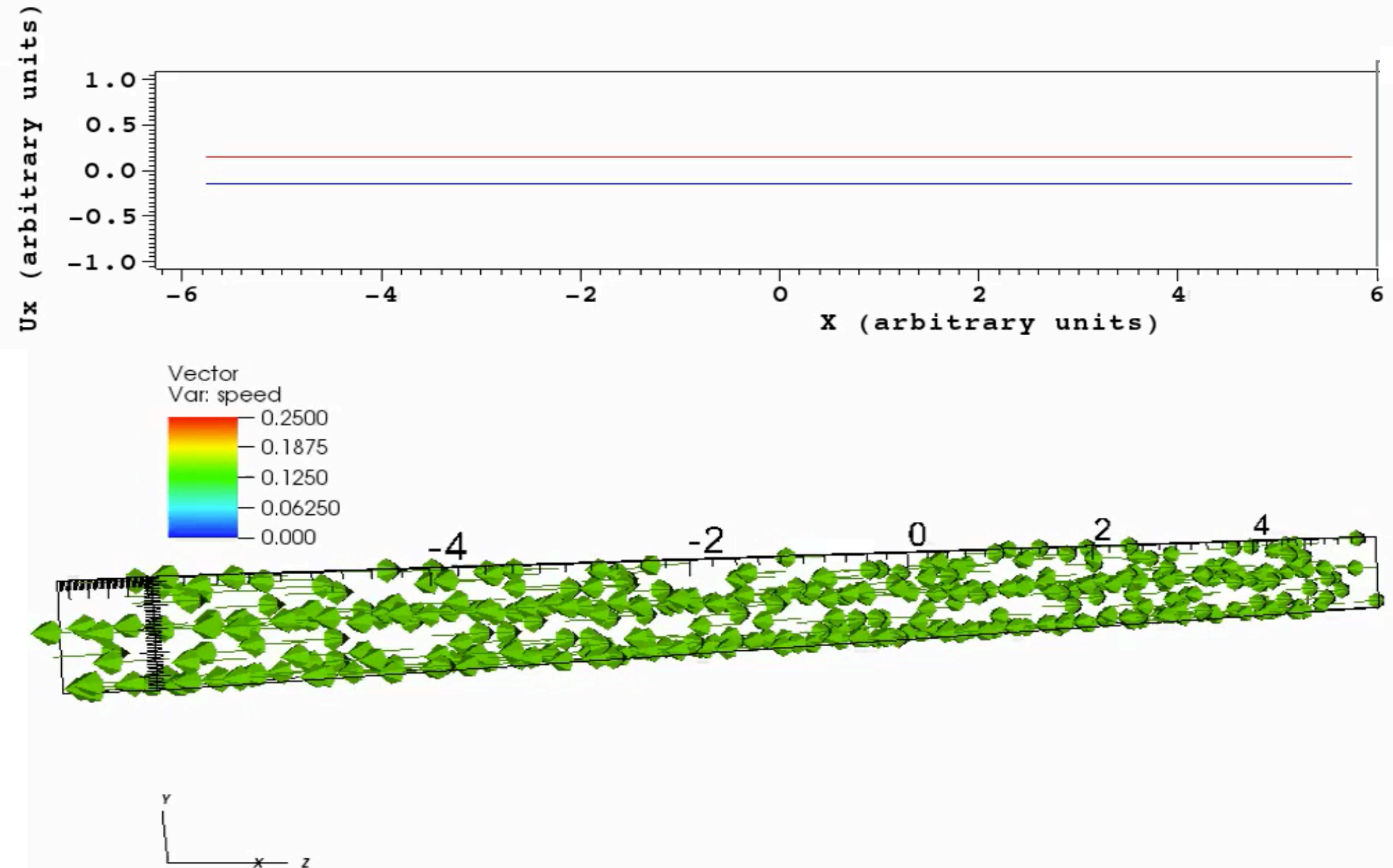
Domain size  
physical 3.0  $R_{LC}$   
with PML 3.6  $R_{LC}$

Resolution (max)  
720<sup>3</sup> cells  
1.5 10<sup>9</sup> particles

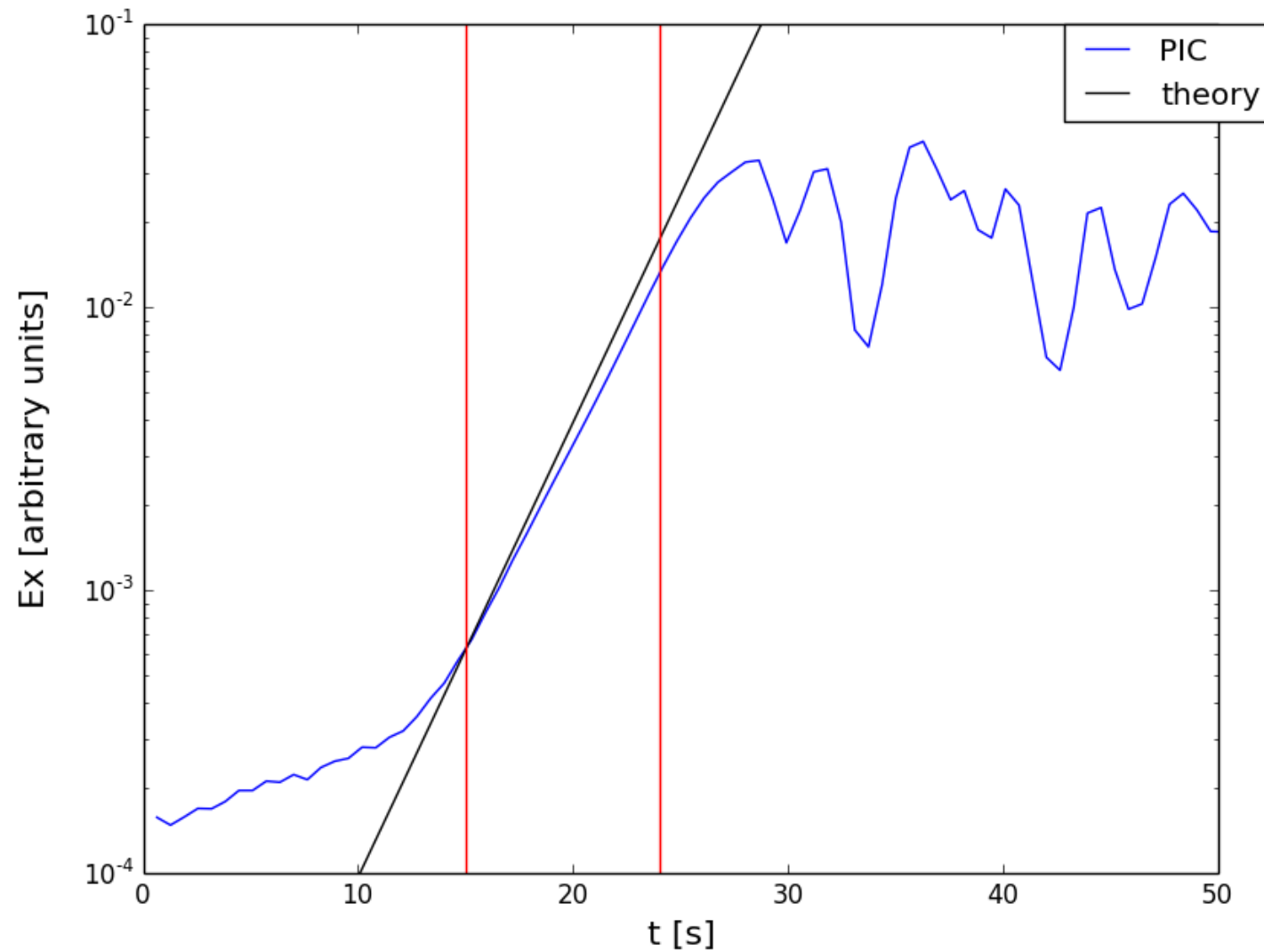
All the simulation run on Discover - NASA NCCS

Most of the plots and videos are done with VisIt (produced by LLNL)

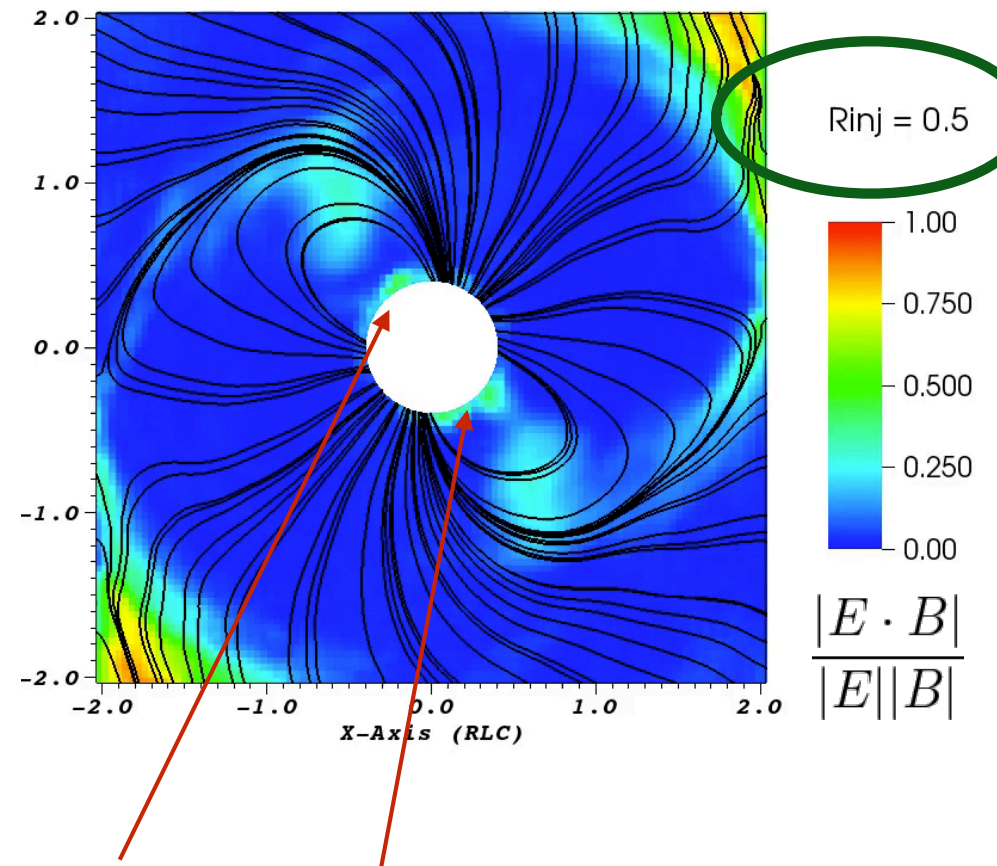
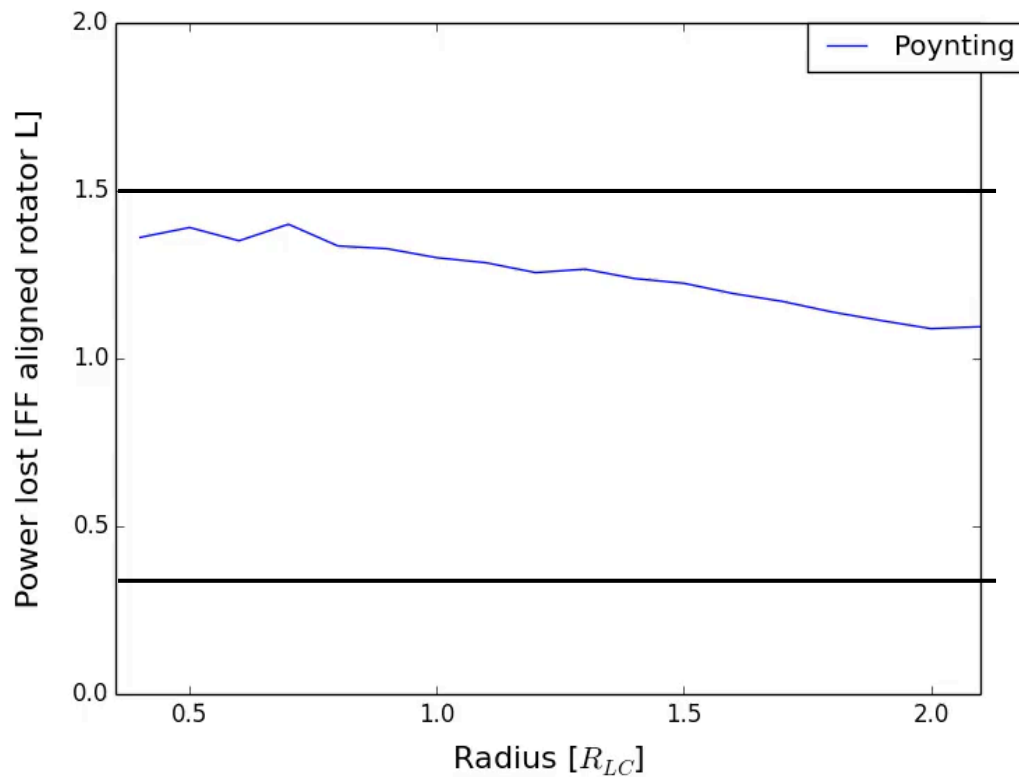
We tested our code with typical plasma problems, like the two stream instability



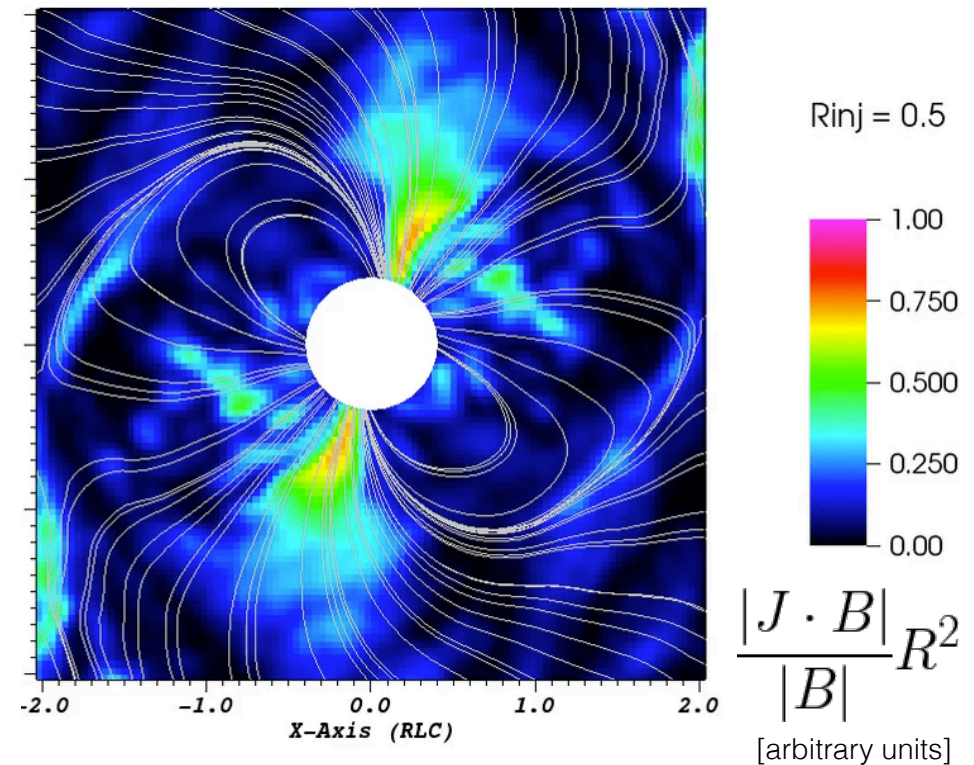
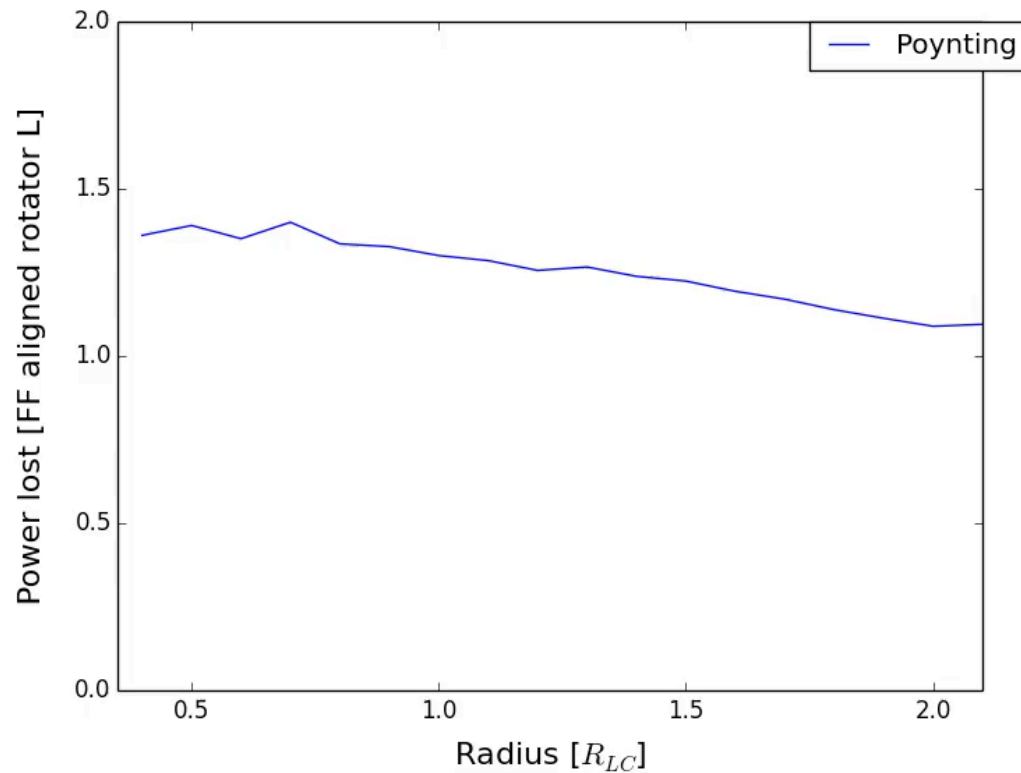
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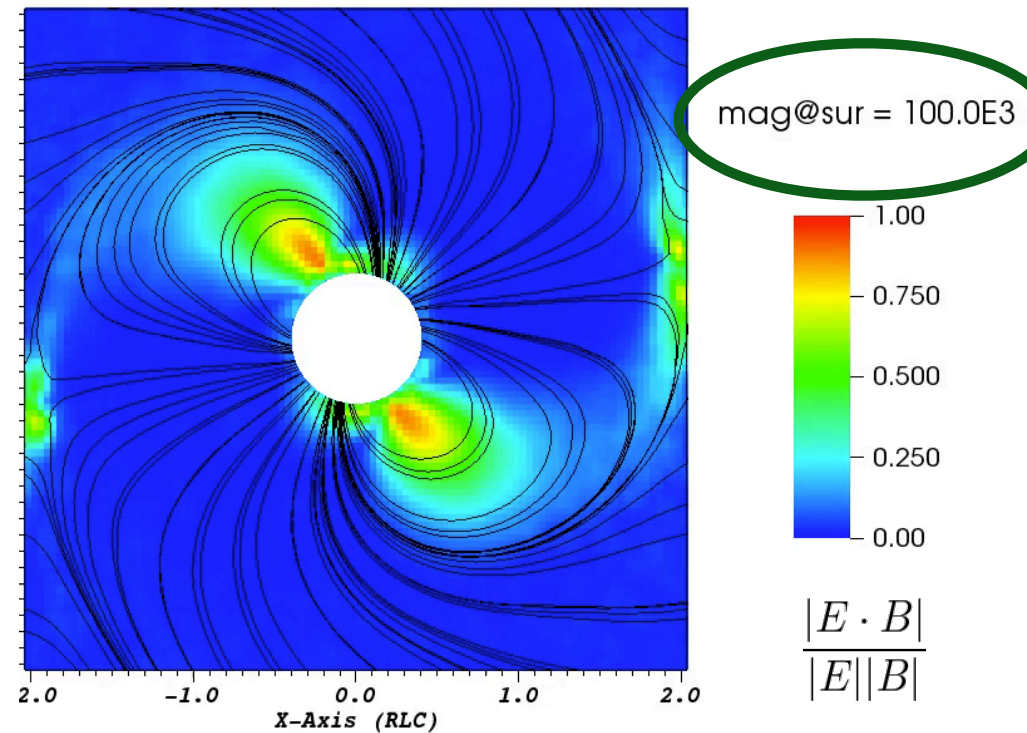
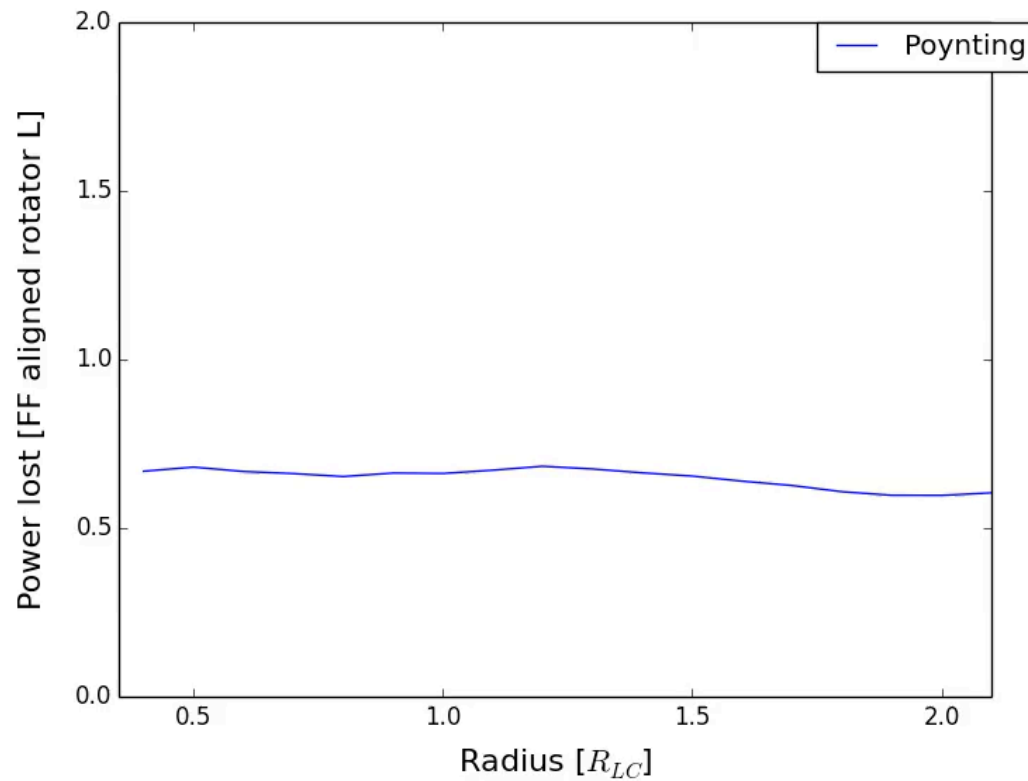
We inject neutral plasma everywhere until we reach a lower magnetization limit inside a fixed radius



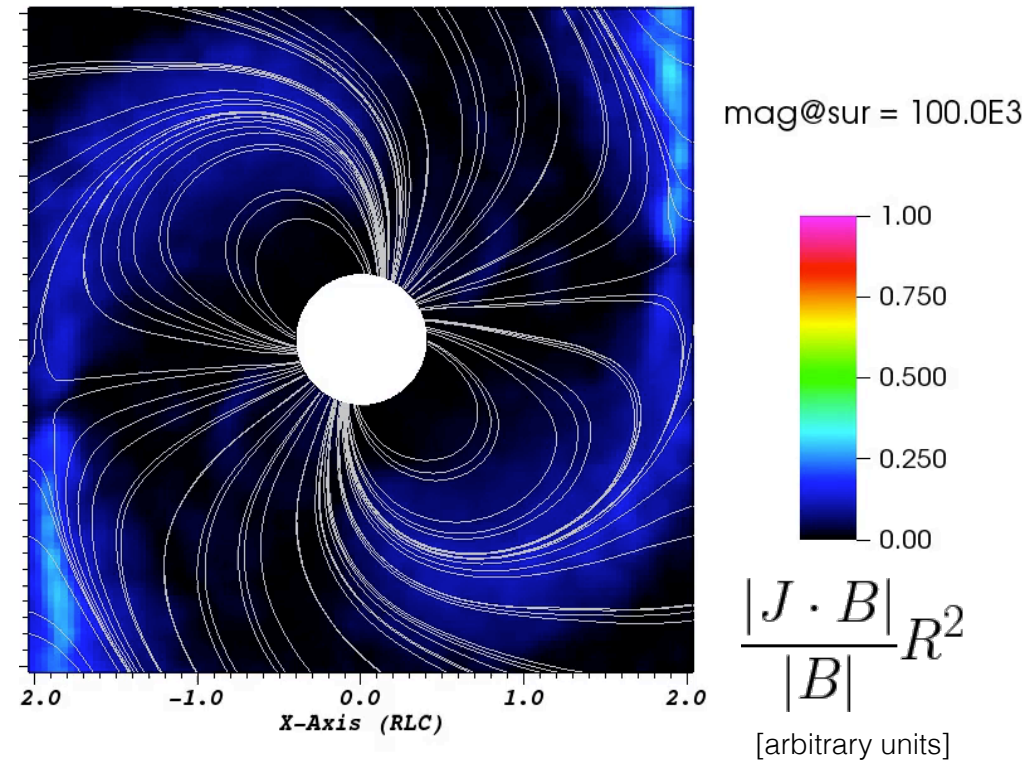
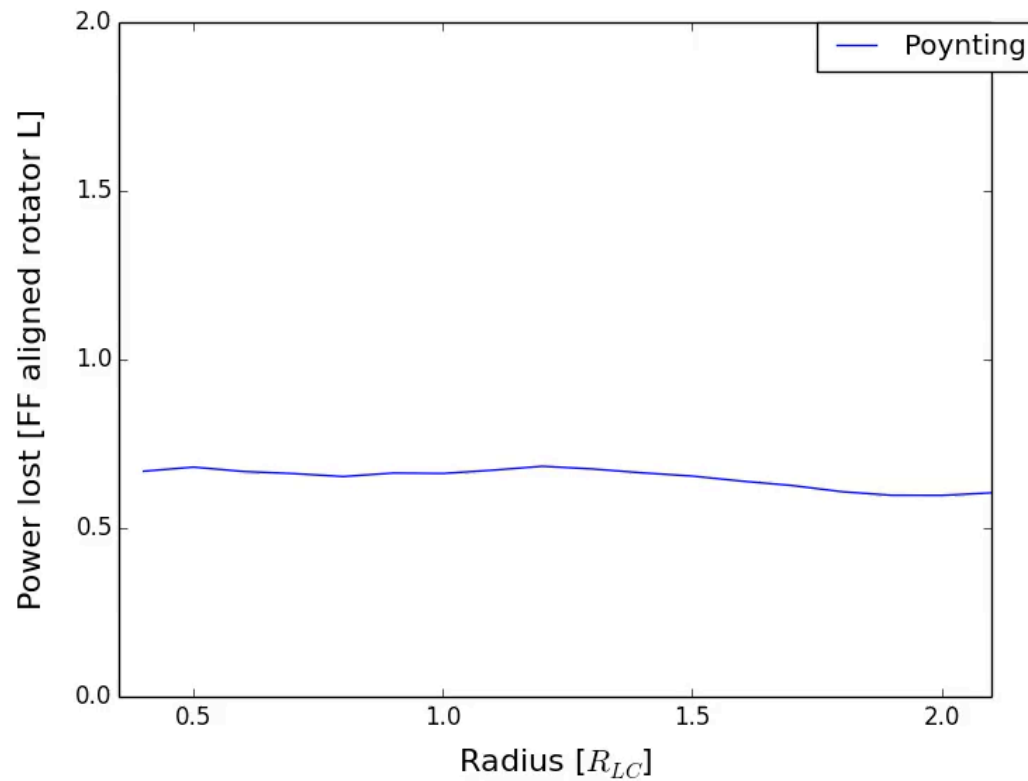
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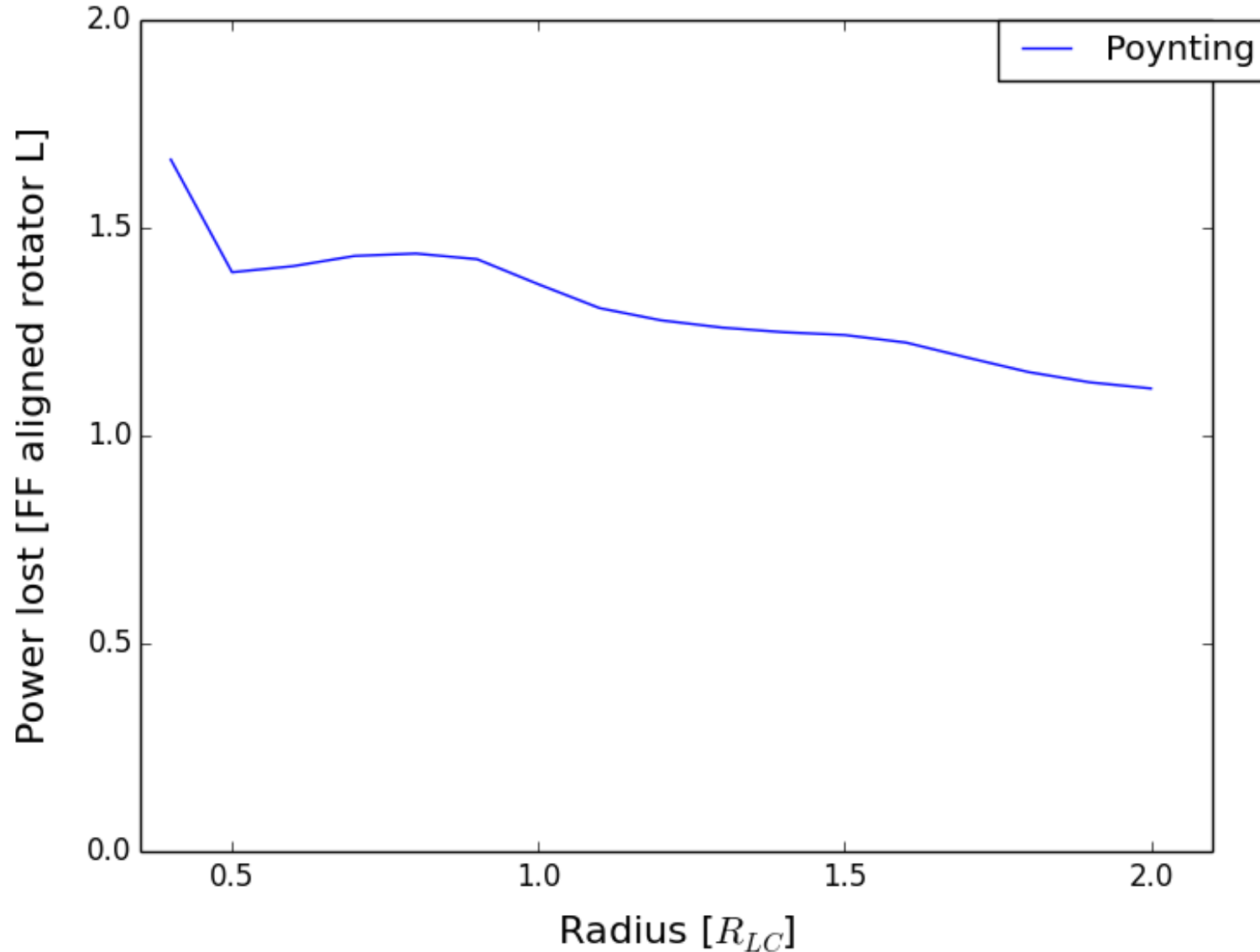
# We inject neutral plasma everywhere until we reach a lower magnetization limit inside a fixed radius



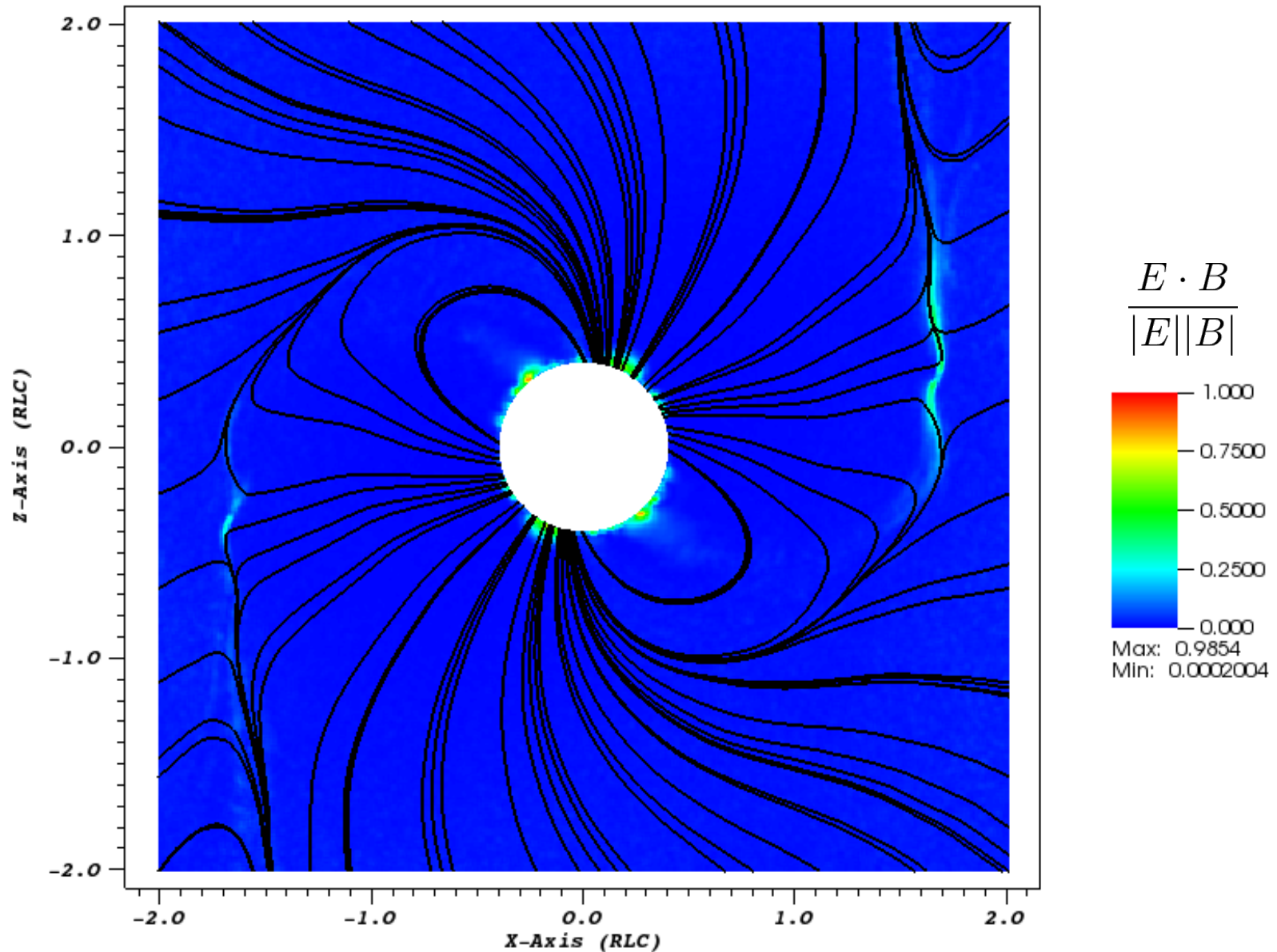
# We inject neutral plasma everywhere until we reach a lower magnetization limit inside a fixed radius



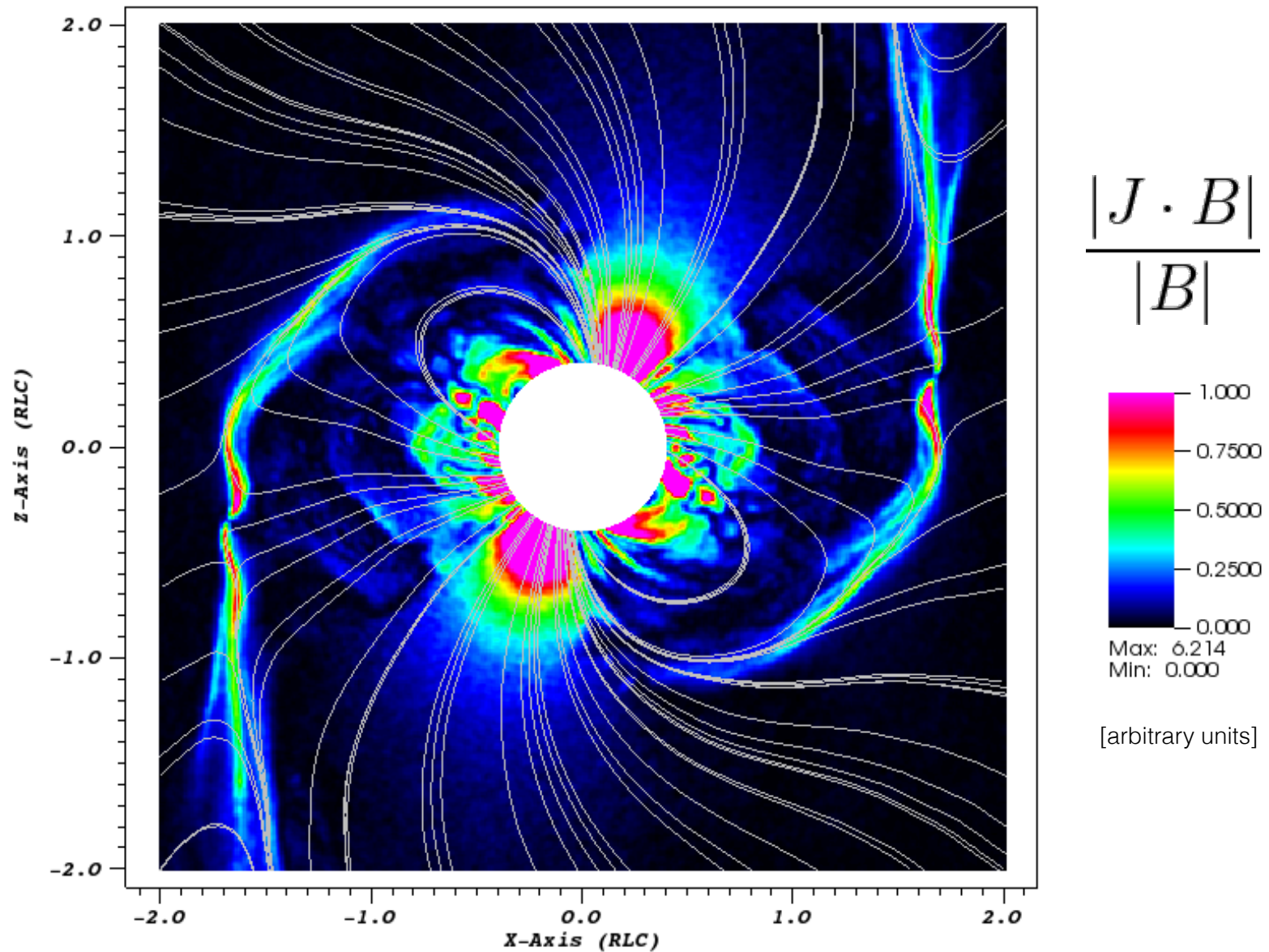
We reached a configuration similar to the force free gradually changing the injection parameters



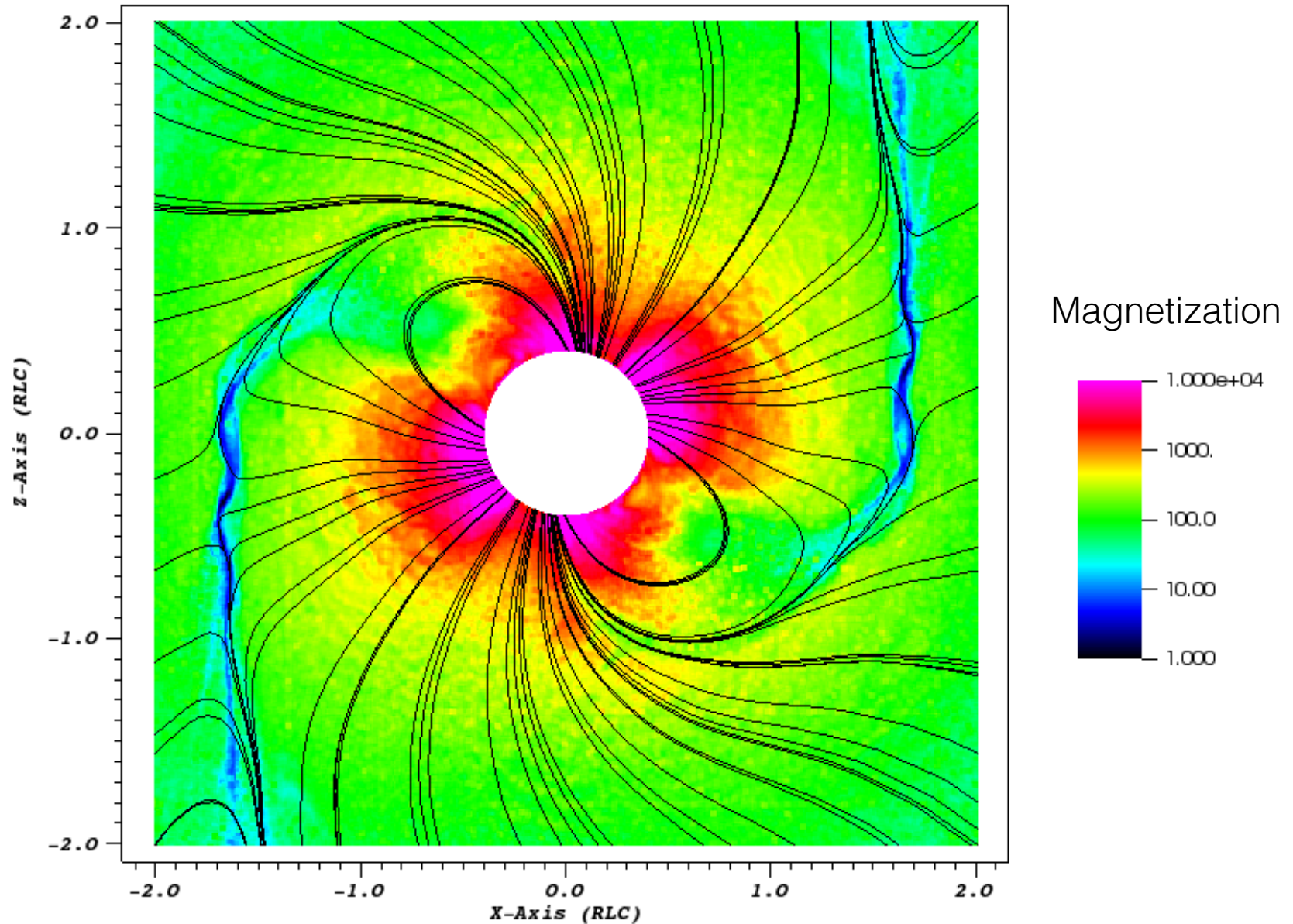
In the force free solution the parallel electric field is screened except for the current sheet



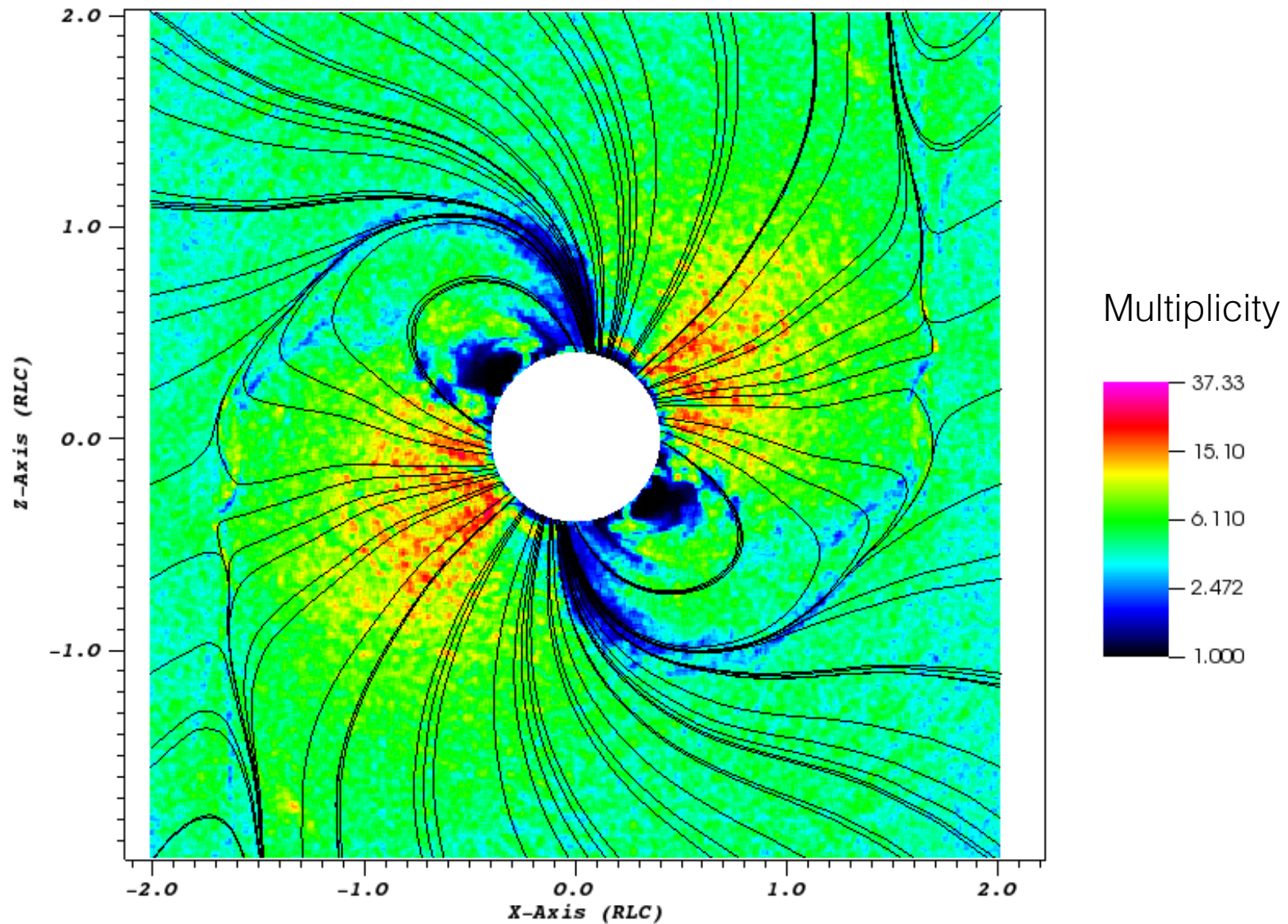
The currents flow on the current sheet and from the polar cap



It is necessary to keep the magnetization high to resemble the real pulsar behavior

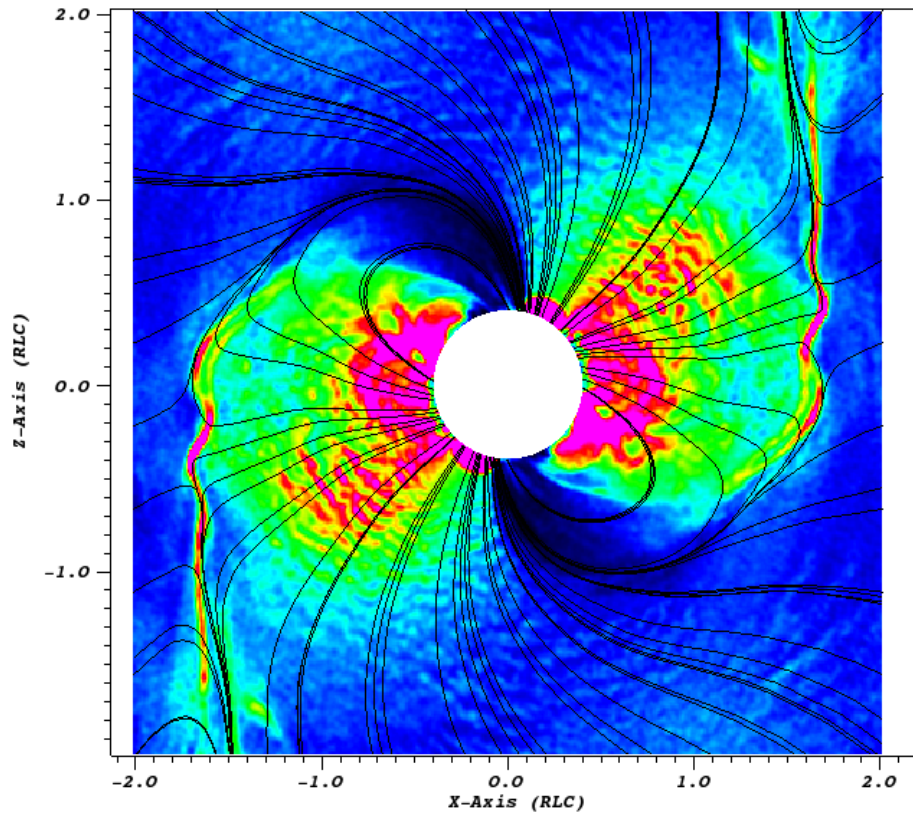


In PIC simulations the multiplicity is lower than in a real pulsar



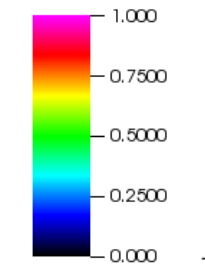


We can look at the different contribution of electrons and positrons to the current

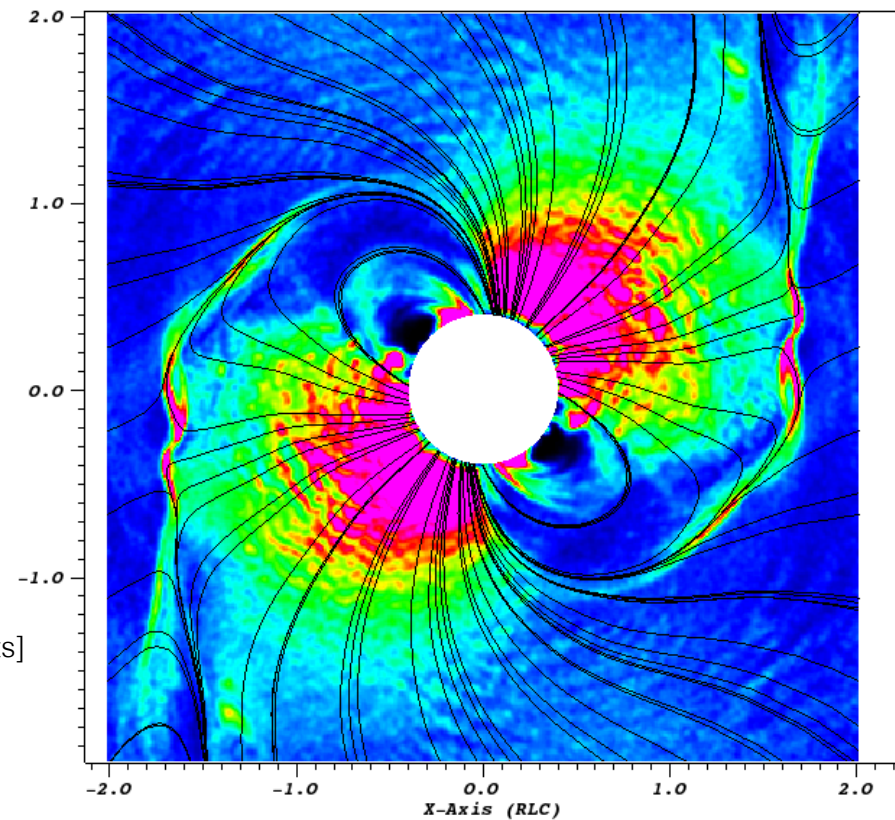


Positrons

$$\frac{|J \cdot B|}{|B|}$$

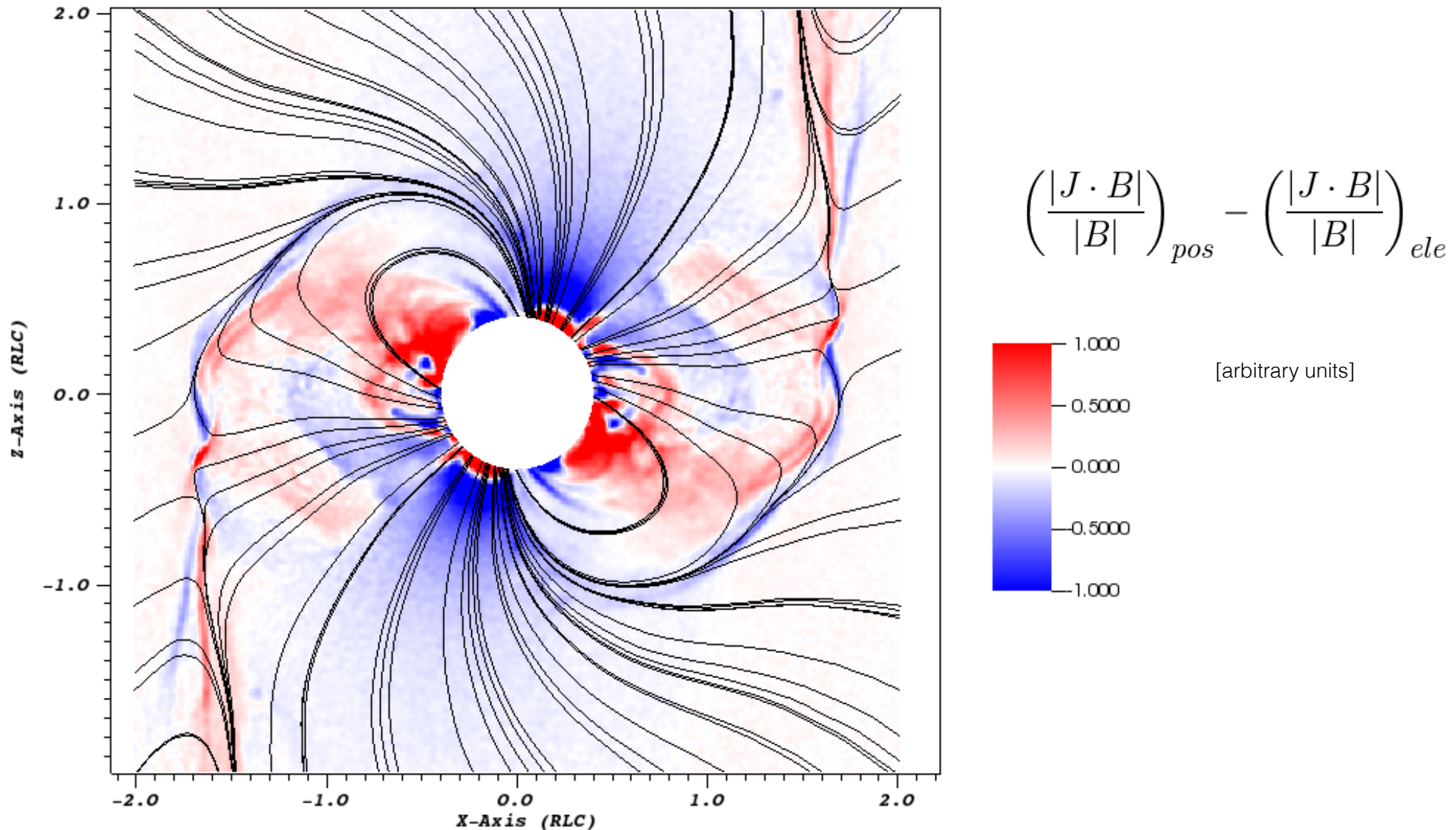


[arbitrary units]

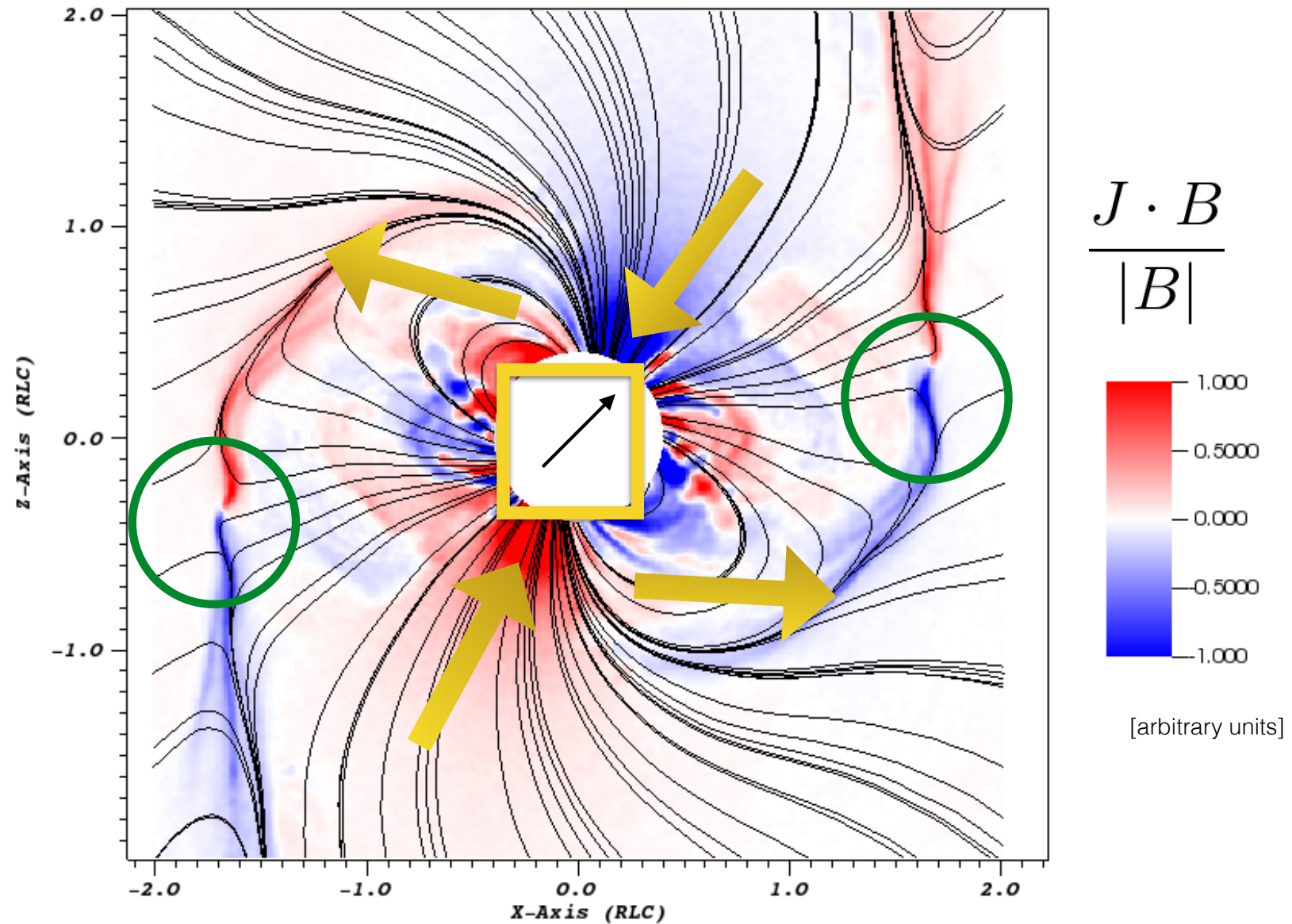


Electrons

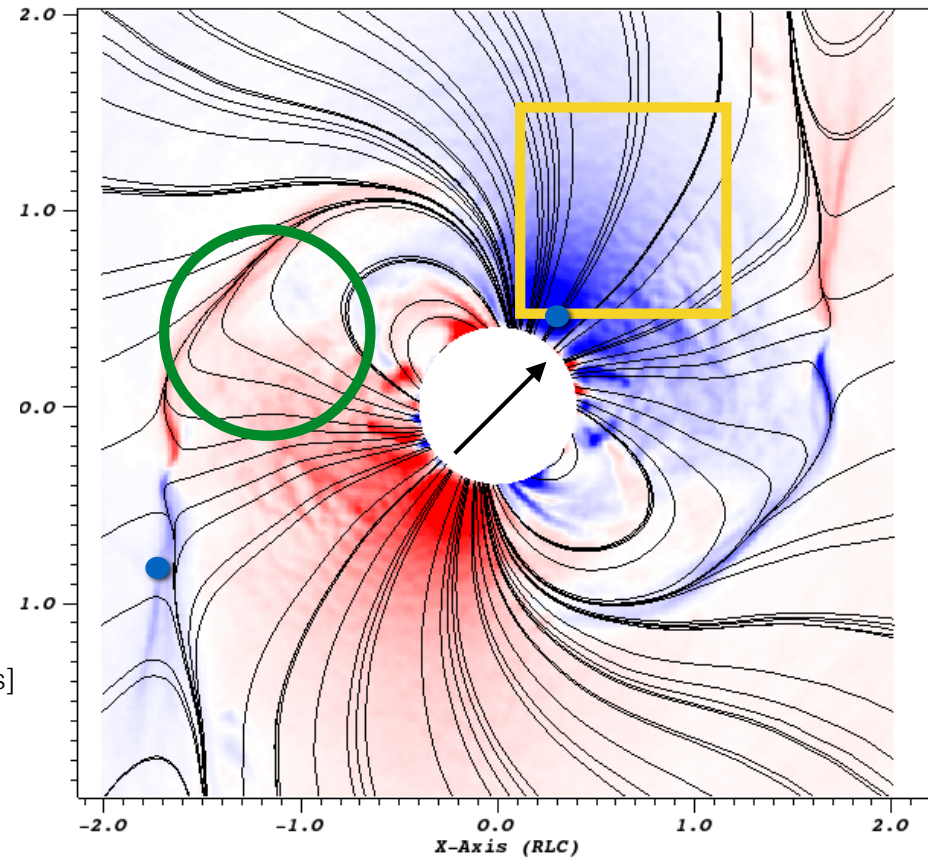
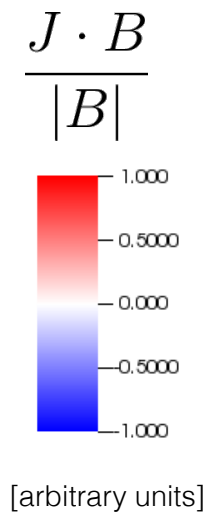
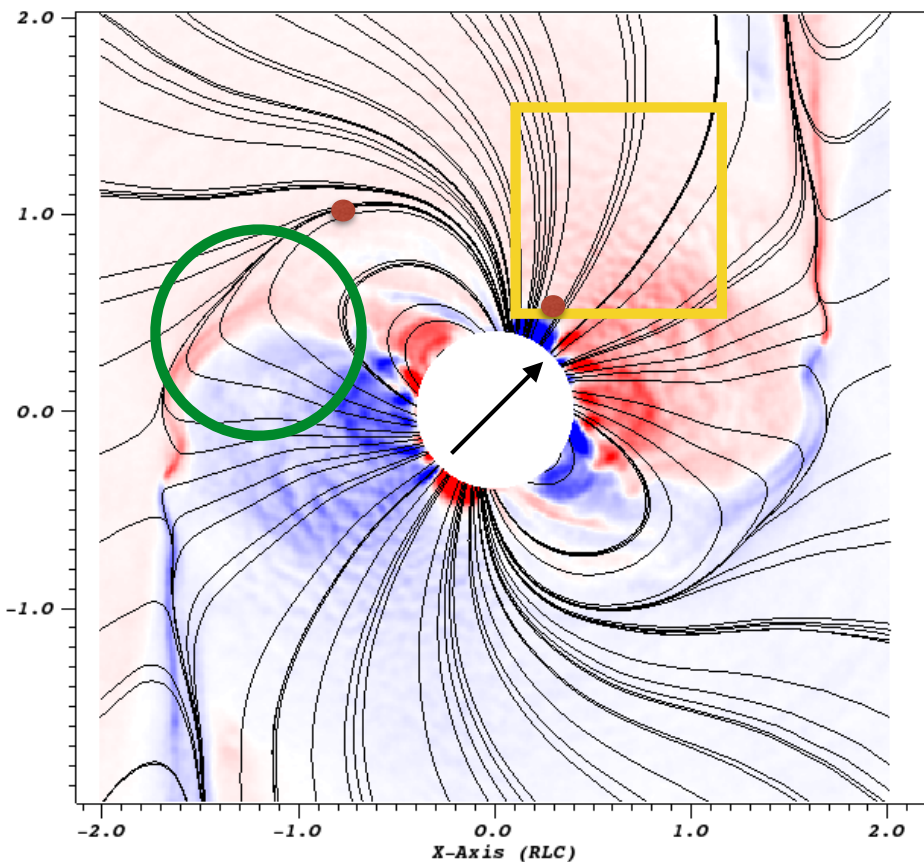
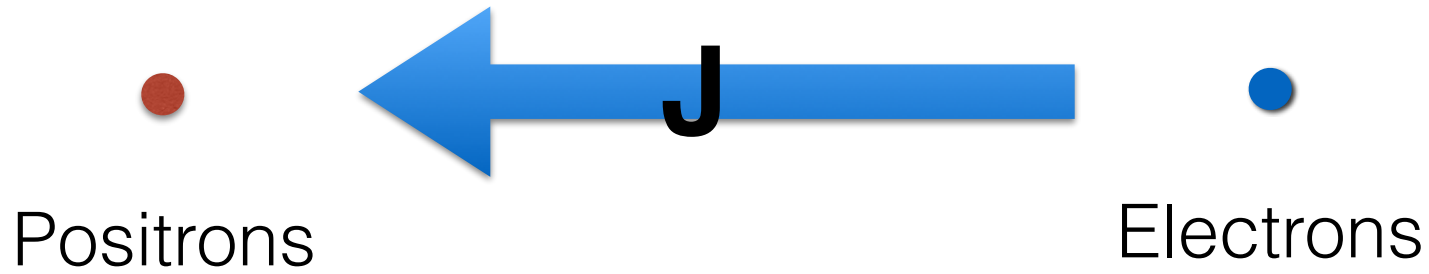
... And we can see where electrons mostly contribute to the current and where positrons do



We can also look at the direction of the flows

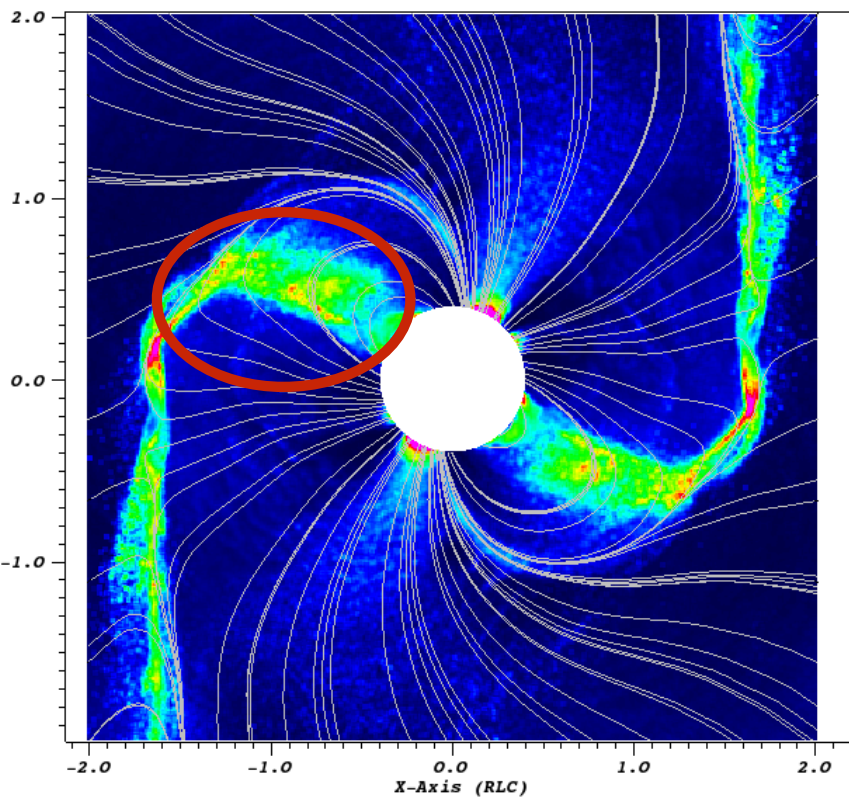


There are zones in which the flows counter stream and others in which they flow in the same direction

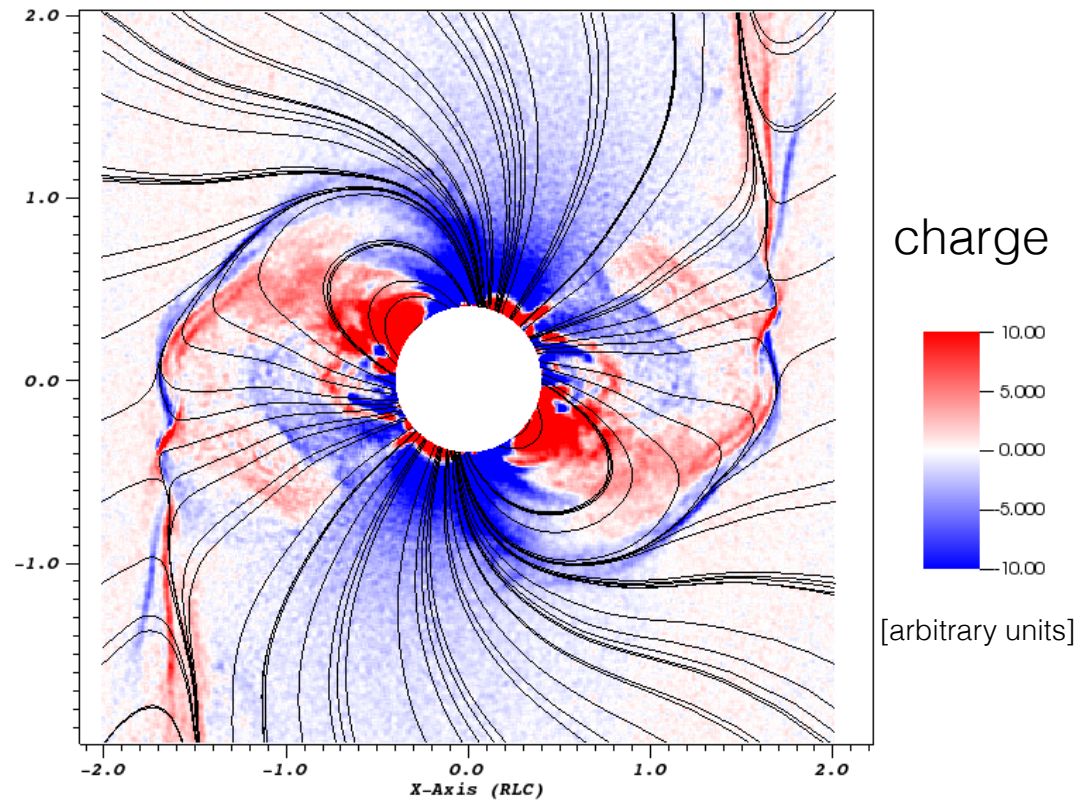
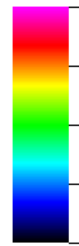


Looking at the energy of the particles we see the most energetic flow in the branches of the current sheet

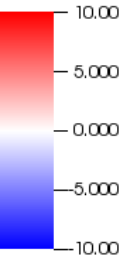
$$\left( \frac{|J \cdot B|}{|B|} \right)_{nos} - \left( \frac{|J \cdot B|}{|B|} \right)_{el\rho} \quad [\text{arbitrary units}]$$



$\gamma$

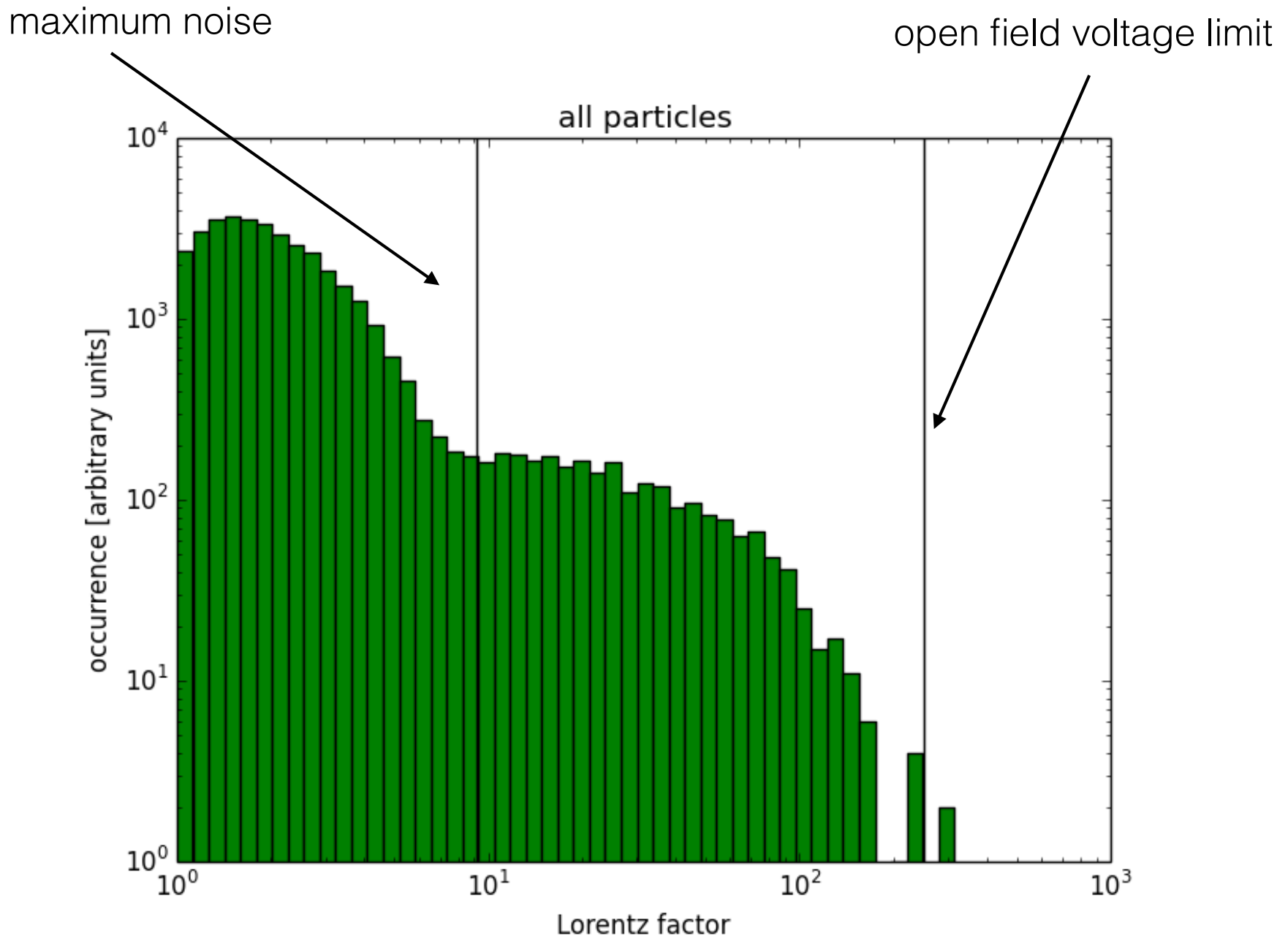


charge



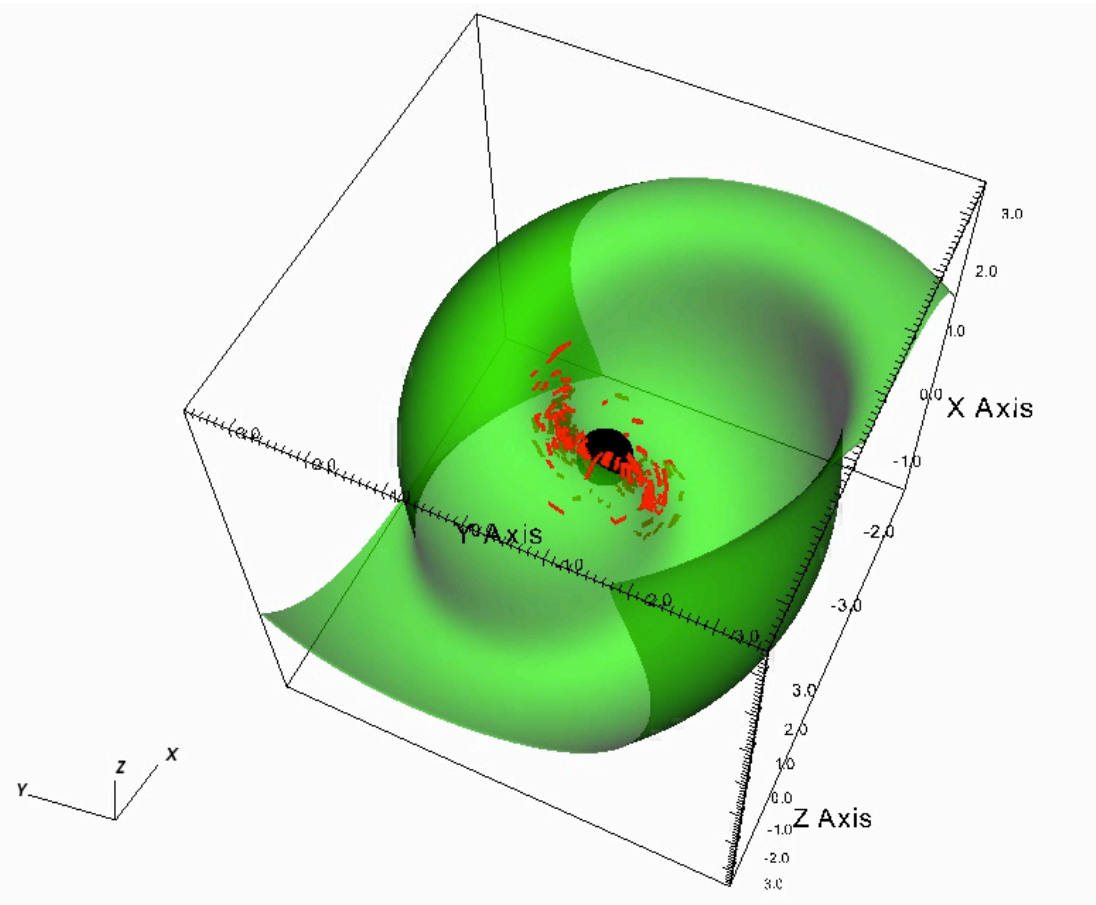
[arbitrary units]

# The knowledge of the particle energy distribution is limited by the noise and the magnetic field



**In the whole 3D structure the most energetic particles flow out on the current sheet**

corotating frame



inertial frame

